### Critical phenomena: A semi-linear wave model

Isabel Suárez Fernández



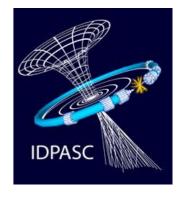


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#### In collaboration with: David Hilditch & Rodrigo Vicente

### Motivation

#### Weak Cosmic Censorship Conjecture

Generically, there is global existance outside a black hole.

There is no hope, right now, to prove it in general for General Relativity.

#### Counterexamples

#### Critical phenomena in General Relativity

Discovered by M. Choptuik in 1993 [1] in simulations of spherically symmetric collapse of a scalar field.

We consider one-parameter families of initial data that takes a critical value *a*\* and divides the solutions space between the

Solutions near the critical point show critical behaviour: Universality, power-law mass scaling of the BH and self-similar be

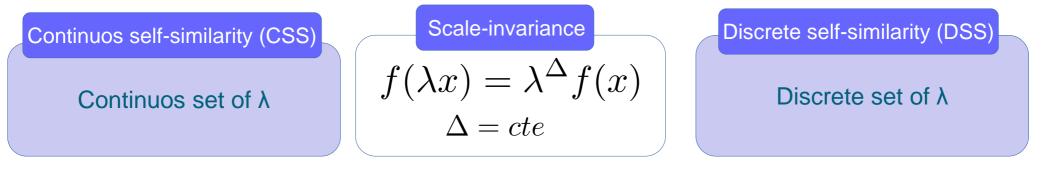
#### Self-Similarity of the critical solution

Type of scale-invariant symmetry.

The critical solution in GR shows Discrete Self-Similarity.

There is no known simple model that has DSS threshold solutions [2].

### Self-similarity and scale-invariance

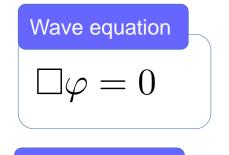


What is the structural difference behind the models?

Seashell as example of CSS (allowing rotations)

Koch curve as example of infinite DSS

### Semi-linear wave equations



Deformation

$$\phi = D[\varphi]$$

Solution in one dimension 
$$\varphi = \frac{f(t+r) - f(t-r)}{r}$$
 Working

 $\Box arphi(\phi) = 0$ 

What do we need? •A solution to the wave equation •A deformation function •Some mathematical work •We work in one dimension

Semilinear wave equation

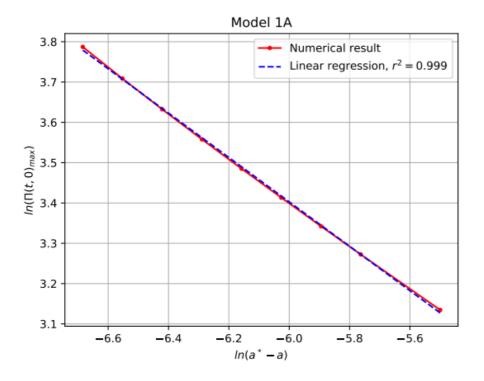
$$\Box \phi - \frac{D'[\varphi]}{D''[\varphi]} \partial_{\alpha} \phi \partial^{\alpha} \phi = 0$$

What do we get? .Semilinear wave equation .Analytical solution .Self-similar solutions .Same structure as GR

### Model 1 – Results

#### The model 1

$$\Box \phi - \partial_{\alpha} \phi \partial^{\alpha} \phi = 0$$



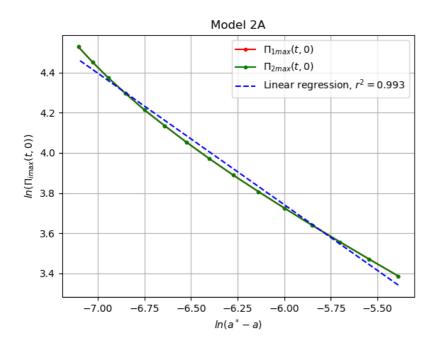
# Analytical solution $\phi = -\ln(\varphi+1)$

What do we get? •Analytical solution to a non-linear •equation •Global solutions for small data •CSS threshold solutions

### Model 2 - Universality

#### The model 2

$$\Box \phi_1 - \partial_\alpha \phi_2 \partial^\alpha \phi_2 = 0$$
$$\Box \phi_2 - \partial_\alpha \phi_1 \partial^\alpha \phi_1 = 0$$



#### What do we get?

.This model does not come from deformation of the wave e

.Global solutions for small data

.CSS threshold solutions

Universal behaviour near thecritical point

•Solutions of model 1 are embedded in its solution space.

.The threshold solution always seems to correspond to the

### Towards DSS

## Deformation $\phi = D[\varphi]$

Working...
$$\Box arphi(\phi) = 0$$

#### What do we need?

In a DSS solution it is the **<u>derivative</u>** of the field what blows up.

Our massless scalar field itself has to be **bounded**.

Careful definition of the deformation function

Defining

$$D[\varphi] \equiv P \circ C[\varphi]$$

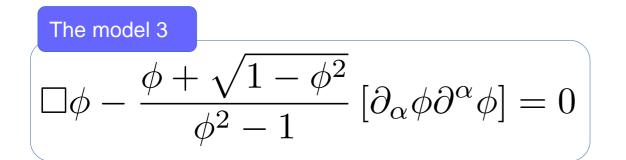
#### Where

**P**: Bounded periodic function of period  $\Delta$ .

**C**: Compactifying function.

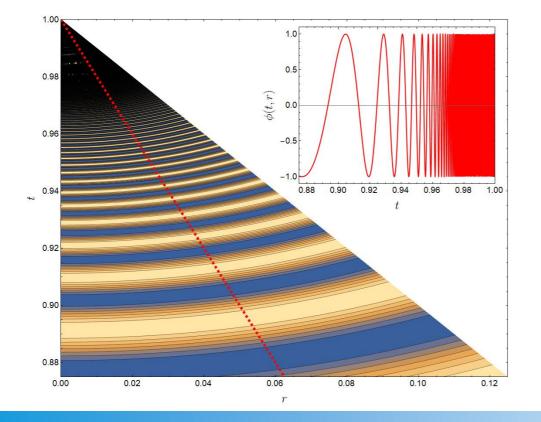
### Will we have DSS now?

### Model 3 – Discrete self-similarity



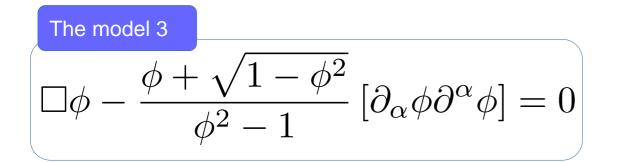
Analytical solution

$$\phi = \sin(\ln(\varphi + 1))$$



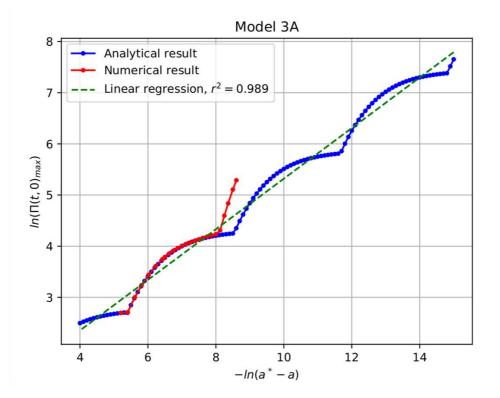
#### We do have DSS!

### Model 3 – Discrete self-similarity



Analytical solution

$$\phi = \sin(\ln(\varphi + 1))$$



### Conclussions & work in progress

#### Tool for models

We have a tool to obtain non-linear models with known analytical solution and critical phenomena for a massless scalar fi

#### First simple DSS model

We constructed the first ever simple model with DSS threshold (analytical) solutions.

We have a tool (code) to test models for which we dont know the analytical solution.

#### Work in progress

We are currently working in modifications of the presented models where the analytical solution is unkown but we can have

The aim is to fully understand the structure that makes a scale-invariant solution CSS or DSS.

### Thanks for your attention

### **Questions?**