

# MARTA readout system

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## Motivation

- Build a muon detector for cosmic ray ground array experiments
- Placing Resistive Plate Chambers (RPC) detectors under Auger's water Cherenkov detectors (WCD)
- Help in the understanding of the muon number discrepancies between data and models





## Challenges

- Instrument the RPCs:
  - Widely used gaseous detectors of charged particles
  - Avalanche mode: Fast pulses that need pre-amplification
  - Estimate the number of muons using both a simple threshold measurement and a charge measurement
- Comply with the strict demands of outdoor field operation:
  - Limited power and space
  - Harsh weather conditions
  - Low maintenance



### Design and characterization of the PREC (Prototype Readout Electronics for Counting particles)

2920

IEEE TRANSACTIONS ON NUCLEAR SCIENCE, VOL. 65, NO. 12, DECEMBER 2018

The MARTA (Muon Array With RPCs for Tagging Air Showers) Front-End Acquisition System

#### Charge distribution measurement of MARTA's RPCs

RPC hodoscope for SSD quality control

GAP-2019-017

#### doi:10.1088/1748-0221/11/08/T08004

 Generic DAQ: used as a proof of concept without complying with the operation demands

#### doi:10.1109/TNS.2018.2879089

 Final system to be used: the system is able to read the fast RPC signals while complying with the operation demands.

#### To be submitted

 RPC charge measurement description and analysis

Collaboration internal gap note

• Application example: a test system for the new Auger's scintillators was successfully build



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Design and characterization of the PREC (Prototype Readout Electronics for Counting particles)

- A proof of concept DAQ system was build using analogue electronics:
  - A simple amplifier followed by a comparator discriminate the RPC signals
- In this work a full characterization of the system performance was presented



• The charge gain was characterized for different temperatures and operation voltages



- No crosstalk was found above the noise level of 1.6 mV
- Trigger efficiency and signal propagation time in the board were studied
- Power consumption was found to be of about 1 Watt per channel

- The system was found to be a good fit for RPC signals
- However it occupies to much space and it draws to much power to be used in field operation
- It was used to study the water tank signals that were then compared with simulation
- This study helped validate the collaboration official simulation





in P. Assis et al., Measurement of the water-Cherenkov detector response to inclined muons using an RPC hodoscope, ICRC 2015

- The MARTA (Muon Array With RPCs for Tagging Air Showers) Front-End Acquisition System
  - A second acquisition system was developed
  - It was designed to comply field operation:
    - Low power: 1.43 Watts per 64 channels
    - Compact: 22 x 14 cm<sup>2</sup>
  - Based on the MAROC ASIC: 64 input channels, preamplifier, threshold and charge measurement
  - 2 prototypes and a final version were build
  - In this work a detailed description of how the system works is given and validation tests are performed

- The ASIC is responsible for the digitization of the RPC signals while the FPGA performs the digital electronics: measurement and data management and communications
- Responsible for the development of the firmware, software and test benches used to validate the design





### Validation results



A low power RPC acquisition system was designed, debug and validated.

### **Charge distribution measurement of MARTA's RPCs**

- The charge spectra of the RPC is a direct consequence of the avalanche that is developed inside the detector
- The expected shape of this spectra would be exponential (small avalanches)
- Several models have been proposed describing the avalanche evolution, but no model was able to describe it fully
- This spectra can give us good indications about the RPC behaviour and health



in E. C. Zeballos et al., Pure avalanche mode operation of a 2 mm gap resistive plate chamber, 1997

One of the proposed models used a gamma distribution to describe the spectra

$$P(\mathcal{N}_e) = \frac{e^{\frac{\mathcal{N}_e}{G/r}} \mathcal{N}_e^{m-1} (G/r)^{-m}}{\Gamma(m)}$$

• This function was shown to fit data well



in P. Fonte, Analytical calculation of the charge spectrum generated by ionizing particles in Resistive Plate Chambers at low gas gain, 2013  In this work, we described the measurement, optimize the measurement parameters and compared data obtained with the gamma distribution





### RPC hodoscope for SSD quality control

GAP-2019-017

 We built an hodoscope, using MARTA's RPC unit to test the new scintillators that are being installed in the Auger array



Top view

• A smaller scintillator was used









• Data angular distribution





- The RPCs coincidence triggers a measurement in the test scintillator
- We were able to measure the signal vs the particle position



Average signal peak (mV) Average signal peak (mV) y(cm) y(cm) -10 **RPCs** Scintillator Scintillator's box 80.00 ŀ 0⊢ x(cm) x(cm)

- We were able to build and debug the hodoscope using a test scintillator
- The next step is to make it work with Auger's scintillator and acquisition system
- The new stand was build where 4 RPCs will test the new detector for any broken fibbers or inefficiencies



## Conclusions

- A proof of concept DAQ system based on analogue electronics was characterized
- A DAQ system for avalanche RPCs was build, tested, debugged and will be deploy
- It comply with the strict demands of field operation and is able to amplify and measure the fast RPC signals
- Can perform charge measurements with the spectrum following the expected distribution
- We built a test system for the new auger scintillators

## Next

- Finish writing the presented work
- Write the thesis
- After:
  - Integrate the RPC hodoscope in the scintillators deployment process
  - Some field tests have already been performed, but more data is needed
  - Deploy a small engineering array of 30 RPCs, that will help determine a calibration point for the muonic component of the air showers
  - It will also be used to cross calibrate other detectors in the array e.g. AMIGA
  - This acquisition system is started to being used in other project like the tomography of the geological structure of the Lousal mine

## Thank you













REPÚBLICA PORTUGUESA

### Cosmic Rays and the Pierre Auger Observatory

- When a cosmic ray particle reaches the atmosphere creates a cascade of particles called Extensive Air Shower (EAS).
- There is a correlation between the particles sampled at ground and the primary cosmic particles.
- The Pierre Auger Observatory is a 3000 km<sup>2</sup> cosmic ray detector, located in Argentina.
- There is a need to get a more precise and independent surface array measurement of the EAS' muon content.
- This measurement would allow to:
  - Distinguish FE and proton showers.
  - Understand the discrepancy between the muon models and data.





Primary cosmic ray

Particle cascade