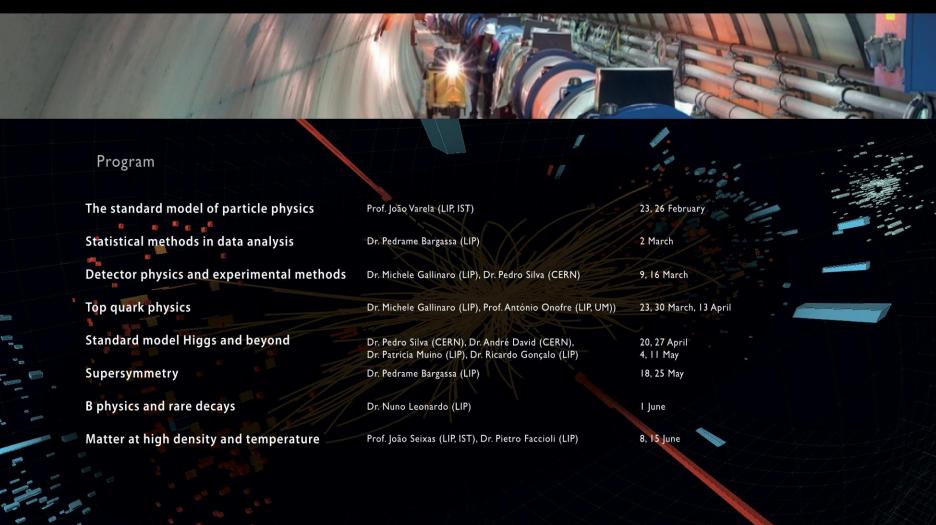


Course on Physics at the LHC

LIP Lisbon, February - June 2015



The lectures will take place on Mondays, between 17:00 and 18:30 at LIP, Av. Elias Garcia 14 r/c, 1000 Lisbon - Portugal



Introduction

Specialized course on the Physics at the Large Hadron Collider organized by LIP in the framework of IDPASC.

The objective of the Course is to introduce the physics, analysis methods and results on the physics of the LHC experiments.

Emphasis is placed on the search for new physics.

Benchmark channels in proton-proton collisions will be discussed in detail:

- identification of the objects involved
- signal and background properties
- background estimation and S/B discriminants
- estimation of systematical errors
- extraction and interpretation of the final results



Program

Program		
The standard model of particle physics	Prof. João Varela (LIP, IST)	23, 26 February
Statistical methods in data analysis	Dr. Pedrame Bargassa (LIP)	2 March
Detector physics and experimental methods	Dr. Michele Gallinaro (LIP), Dr. Pedro Silva (CERN)	9, 16 March
Top quark physics	Dr. Michele Gallinaro (LIP), Prof. António Onofre (LIP, UM))	23, 30 March, 13 April
Standard model Higgs and beyond	Dr. Pedro Silva (CERN), Dr. André David (CERN), Dr. Patricia Muino (LIP), Dr. Ricardo Gonçalo (LIP)	20, 27 April 4, 11 May
Supersymmetry	Dr. Pedrame Bargassa (LIP)	18, 25 May
B physics and rare decays	Dr. Nuno Leonardo (LIP)	l June
Matter at high density and temperature	Prof. João Seixas (LIP, IST), Dr. Pietro Faccioli (LIP)	8, 15 June



Required background

The course is intended for under-graduate or graduate students having basic training in Particle Physics:

Basic concepts

Elementary constituents of matter and interactions. Quantum numbers and conservation rules. Spin and symmetry groups. Relativistic kinematics. Crosssection. Natural units. Mass and lifetime. Resonances.

Structure of matter

Elastic scattering and form factors. Inelastic scattering experiments. Nucleon structure functions. Scale invariance. Quark model. Parton distribution functions. Introduction to QCD.

Fundamental interactions

Introduction to QED. Fermi interaction. Parity violation. Currents V-A and weak doblets. W and Z bosons. Cabibbo angle. Neutral currents. Electroweak interaction. Gauge symmetries. The Higgs mechanism. Weinberg-Salam model. CP violation.



Background bibliography

- F. Halzen and A.D.Martin, 'Quarks and Leptons', John Wiley and Sons (1984)
- D. Griffiths, 'Introduction to Elementary Particles', John Wiley and Sons (1987)
- B.R.Martin, G. Shaw, 'Particle Physics', John Wiley and Sons (1999)



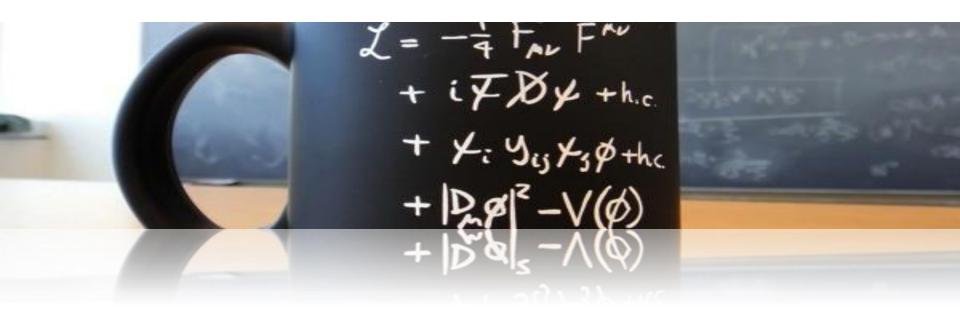
The standard model of particle physics

Lecture1

- 1. The LHC physics case
- 1. The LHC experimental program



The LHC physics case





Particle Physics

Particle physics is a modern name for the centuries old effort to understand the basics laws of physics.

Edward Witten

Aims to answer the two following questions:

What are the elementary constituents of matter?

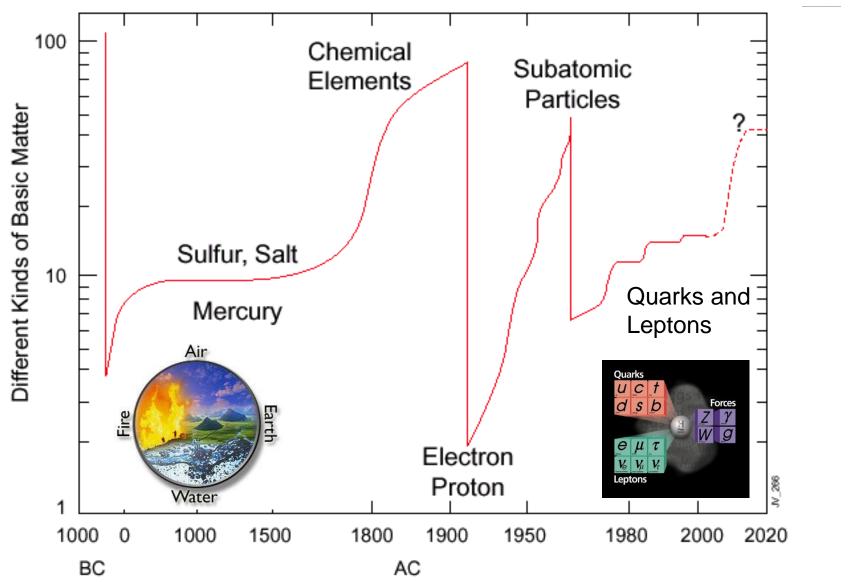
What are the forces that determine their behavior?

Experimentally

Get particles to interact and study what happens



Constituents of matter along History

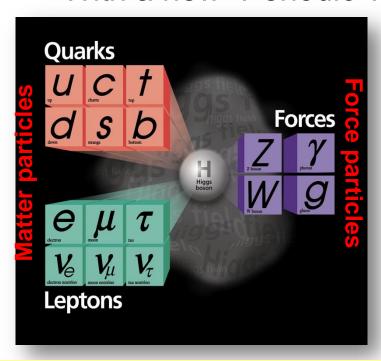




The Standard Model

Over the last ~100 years: The combination of Quantum Field Theory and discovery of many particles has led to

- The Standard Model of Particle Physics
 - With a new "Periodic Table" of fundamental elements



One of the greatest achievements of 20th Century Science

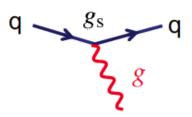
$$L_{H} = \frac{1}{2} (\partial_{\mu} H)^{2} - m_{H}^{2} H^{2} - h \lambda H^{3} - \frac{h}{4} H^{4} + \frac{g^{2}}{4} (W_{\mu}^{+} W^{\mu} + \frac{1}{2 \cos^{2} \theta_{W}} Z_{\mu} Z^{\mu}) (\lambda^{2} + 2\lambda H + H^{2}) + \sum_{l,q,q'} (\frac{m_{l} \bar{l}}{\lambda} \bar{l} l + \frac{m_{q}}{\lambda} \bar{q} q + \frac{m_{q'}}{\lambda} \bar{q}' q') H$$



Standard model interactions

The interaction of gauge bosons with fermions is described by the Standard Model



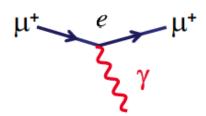


Only quarks
Never changes
flavour

$$\alpha_S \sim 1$$

Gluons massless

EM



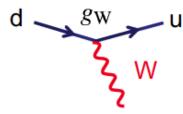
All charged fermions

Never changes flavour

$$\alpha \simeq 1/137$$

Photon massless

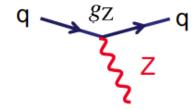
WEAK CC



All fermions

Always changes flavour





All fermions

Never changes flavour

$$\alpha_{W/Z} \sim 1/40$$

W+, Wvery massive **Z**⁰ very massive

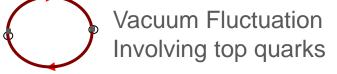


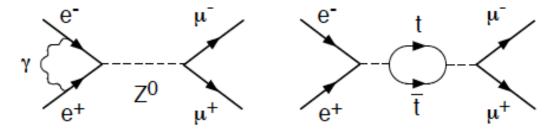
Quantum field theory

A particle-antiparticle pair can pop out of empty space ("the vacuum") and then vanish back into it

These are Virtual particles.

Other examples of Virtual particles:





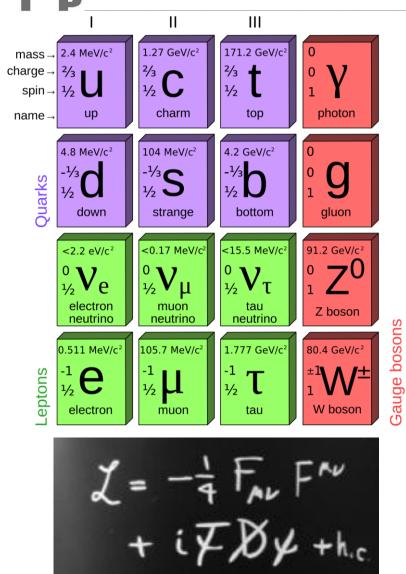
This has far-reaching consequences

The structure of the universe depends on particles that don't exist in the usual sense

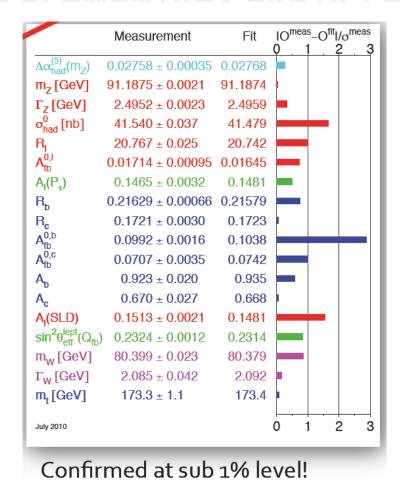
We do not see these particles in everyday life
We must recreate the state of the early hot universe to
make them



SM confirmed by data



STANDARD MODEL OF ELEMENTARY PARTICLES





What's missing?

A "funny" thing happened on the way to the modern theory of quarks, leptons, force fields, and their quanta:

The equations only made sense if all the bosons, and all the quarks and leptons, had no mass and moved at the speed of light!



The Higgs

In the simplest model the interactions are symmetrical and particles do not have mass

The symmetry between the electromagnetic and the week interactions is broken:

- Photon do not have mass
- W, Z do have a mass ~ 80-90 GeV

Higgs mechanism: mass of W and Z results from the interactions with the Higgs field



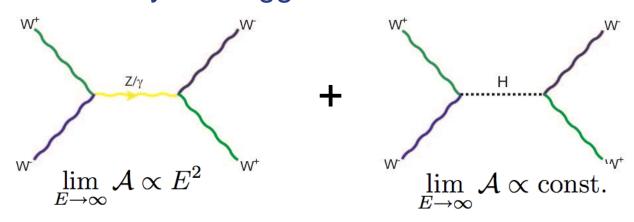
Added bonus

Non-zero average value of the Higgs field can also give masses to the quarks, electrons and muons – to all point-like particles.

Old theoretical problem affecting the quantum theory of the weak force :

the probability of two W's interacting becomes larger than 1 at high energies (> 1 TeV).

This is solved by the Higgs field!





Beyond the standard model

The Standard Model answers many of the questions about the structure of matter. But the Standard Model is not complete; there are still many unanswered questions.

Why do we observe matter and almost no antimatter if we believe there is a symmetry between the two in the universe?

What is this "dark matter" that we can't see that has visible gravitational effects in the cosmos?

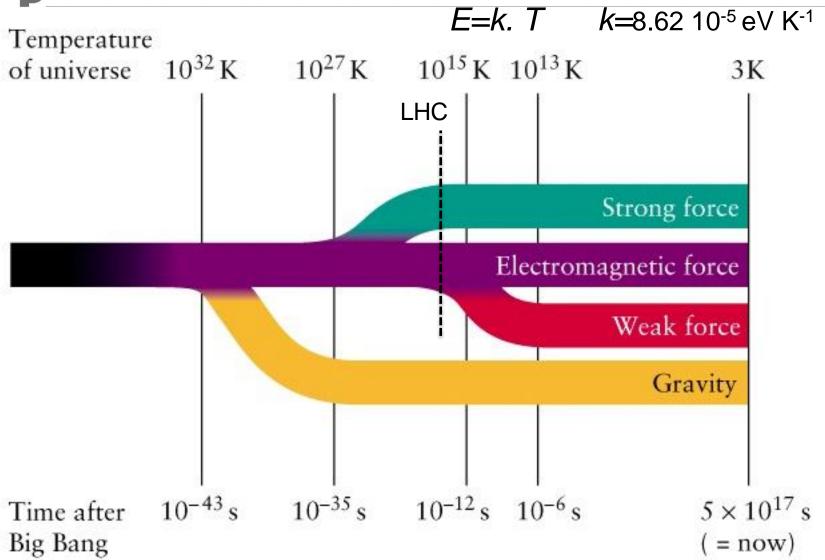
Are quarks and leptons actually fundamental, or made up of even more fundamental particles?

Why are there exactly three generations of quarks and leptons? What is the explanation for the observed pattern for particle masses?

How does gravity fit into all of this?



Forces and expansion of the Universe



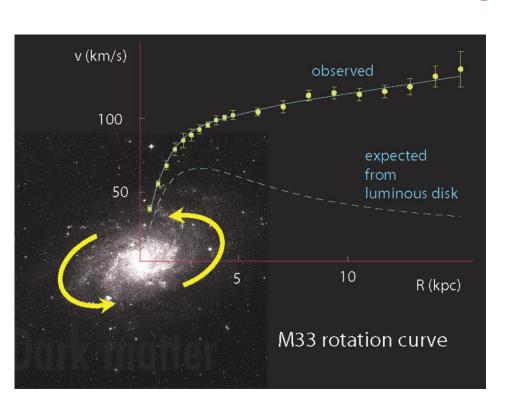


The dark side of the Universe

Long standing problem:

We know that ordinary matter is only ~4% of the matterenergy in the Universe.

What is the remaining 96%?



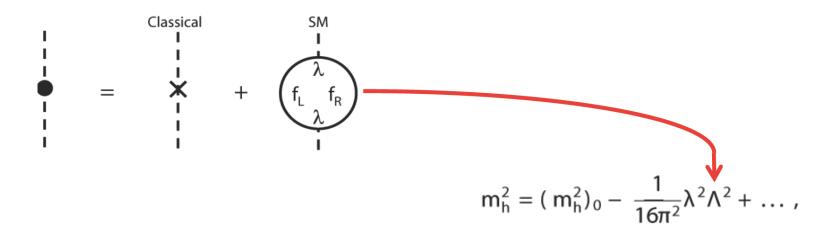
The LHC may help to solve this problem, discovering dark matter



Higgs and hierarchy problem

In the SM the Higgs mass is a huge problem:

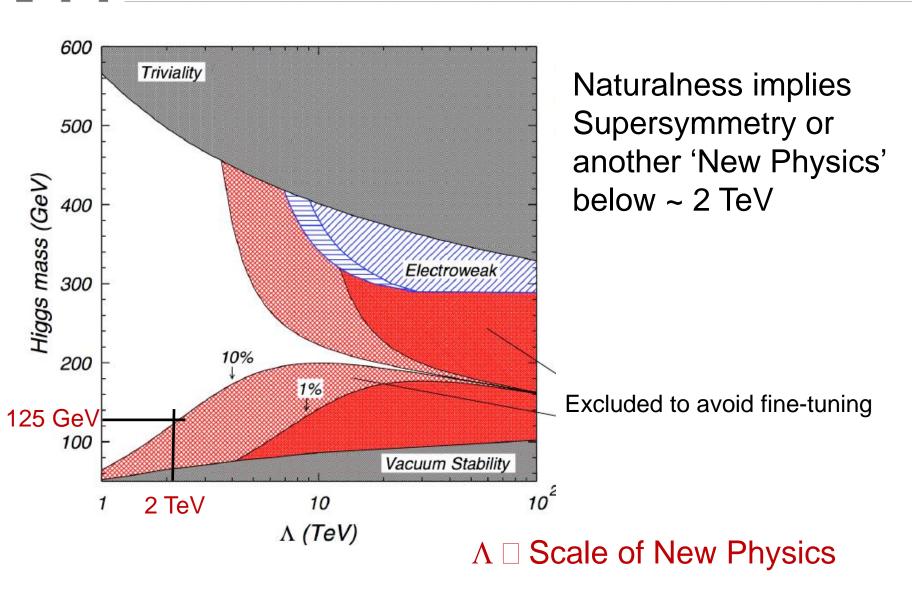
- Virtual particles in quantum loops contribute to the Higgs mass
- Contributions grow with ∧ (upper scale of validity of the SM)
- A could be huge e.g. the Plank scale (10¹⁹ GeV)
- Miraculous cancelations are needed to keep the Higgs mass < 1 TeV



This is known as the hierarchy problem



New physics at a few TeV?





Many possible theories

There are a large number of models which predict new physics at the TeV scale accessible at the LHC:

- Supersymmetry (SUSY)
- Extra dimensions
- Extended Higgs Sector e.g. in SUSY Models
- Grand Unified Theories (SU(5), O(10), E6, ...)
- Leptoquarks
- New Heavy Gauge Bosons
- Technicolour
- Compositeness

Any of this could still be found at the LHC



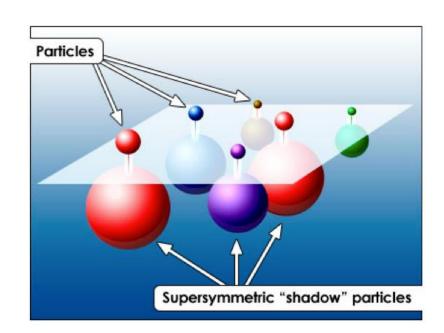
Supersymmetry

Some physicists attempting to unify gravity with the other fundamental forces have proposed a new fundamental symmetry:

- Every fermion should have a massive "shadow" boson
- and boson should have a massive "shadow" fermion.

This relationship between fermions and bosons is called supersymmetry (SUSY)

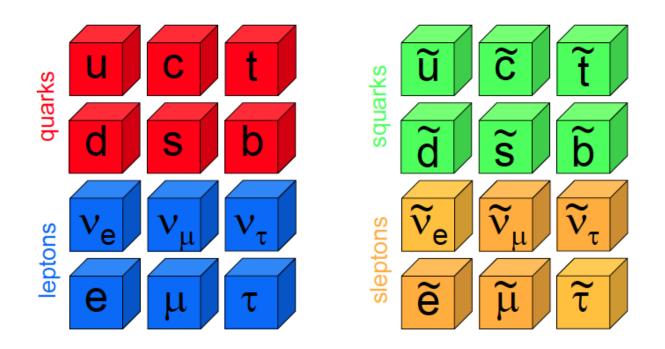
No supersymmetric particle has yet been found, but experiments are underway at CERN to detect supersymmetric partner particles.





Supersymmetry

Double the whole table with a new type of matter?



Heavy versions of every quark and lepton Supersymmetry is broken



Could DM be SUSY particles?

For every "normal" force quanta (boson), there are supersymmetric partners:

photon photino

W, Z bosons Wino, Zino

gluon gluino

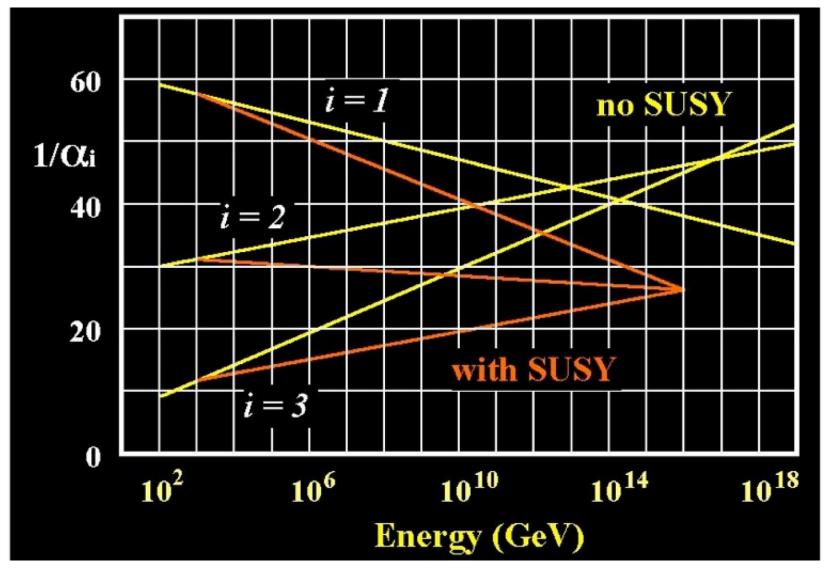
Higgs boson higgsino

These "...inos" are prime suspects to be the galactic dark matter!

Relics from the Big Bang!



The temptation unification





SUSY and the Higgs mass

$$m_h^2 = (m_h^2)_0 - \frac{1}{16\pi^2}\lambda^2\Lambda^2 + \dots,$$

Higgs mass:

- correction has quadratic divergence!
 - $-\Lambda$ a cut-off scale e.g. Planck scale

$m_h^2 = (m_h^2) \left(-\frac{1}{16\pi^2} \lambda^2 \Lambda^2 + \frac{1}{16\pi^2} \lambda^2 \Lambda^2 + \dots \right)$ $\approx (m_h^2)_0 + \frac{1}{16\pi^2} (m_f^2 - m_f^2) \ln(\Lambda / m_f^2),$

Superpartners fix this:

- Need superpartners at mass ~1-2 TeV
 - Otherwise the logarithmic term becomes too large, which would require more fine-tuning.

Cancellation



Extra dimensions

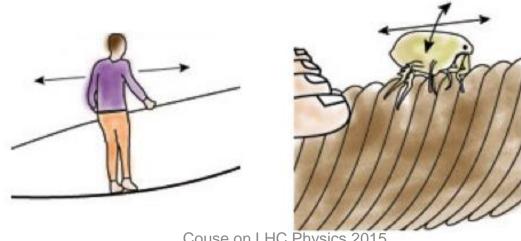
Space-time could have more than three space dimensions. The extra dimensions could be very small and undetected until now.

How can there be extra, smaller dimensions?

The acrobat can move forward and backward along the rope: **one** dimension

The flea can move forward and backward as well as side to side: two dimensions

But one of these dimensions is a small closed loop.





Naturalness

At a Crossroad

Standard Model

 M_{DL}

Multiverse

 $m_h^2 = (m_h^2)_0 - \frac{1}{16\pi^2} \lambda^2 \Lambda^2 + \dots,$

Naturalness

Large Extension Supersymmetry

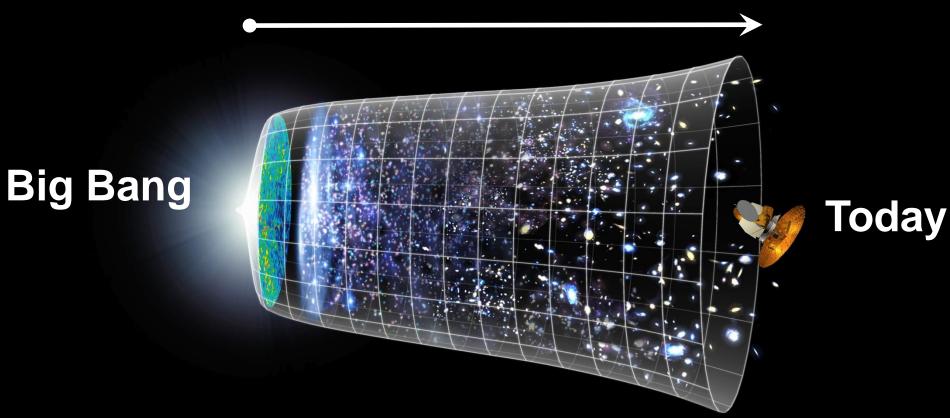
Technicolor

$$\begin{split} m_h^2 &= (m_h^2)_0 - \frac{1}{16\pi^2} \lambda^2 \Lambda^2 + \frac{1}{16\pi^2} \lambda^2 \Lambda^2 + \dots \\ &\approx (m_h^2)_0 + \frac{1}{16\pi^2} (m_{\tilde{f}}^2 - m_f^2) \ln(\Lambda / m_{\tilde{f}}) , \end{split}$$

Savas Dimopoulos, CERN Colloquium, Sep 20, 2012

Timeline of the Universe

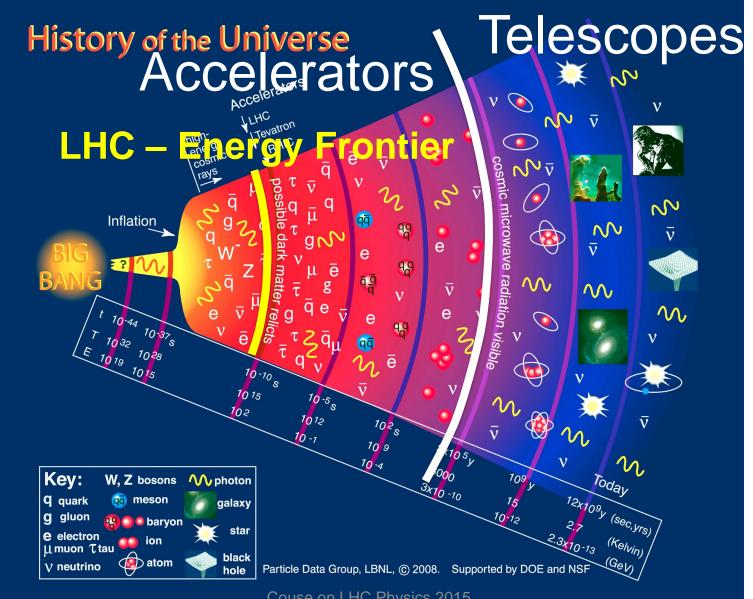
13.7 billion years



LHC recreates the conditions one billionth of a second after Big Bang



Understanding the Universe



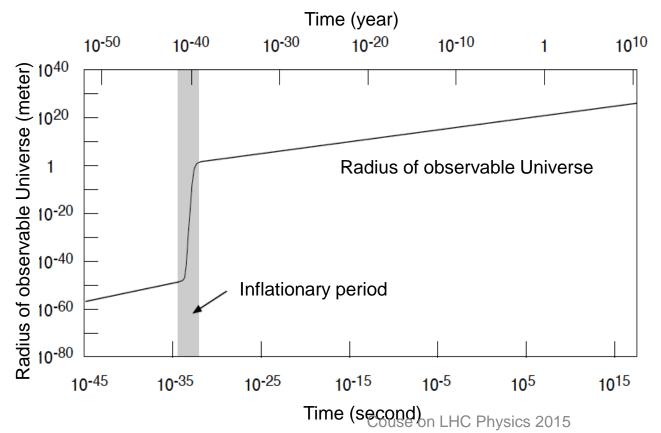


Cosmological inflation

In the very early universe, the physical vacuum undergoes a transition from a high energy state to a low energy state.

The resulting energy shift drives a dramatic exponential expansion.

Explains why the Universe has a uniform Temperature (3 K)



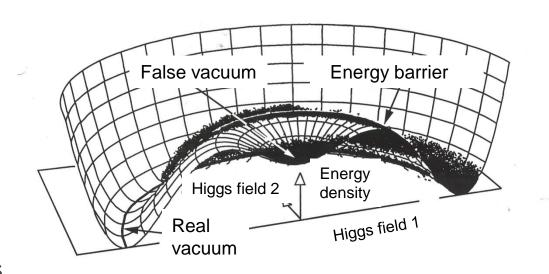
The inflation theory was developed independently in the late 1970's by Alan Guth, Alexey Starobinsky, and others



Higgs like field and inflation

Before the inflation (10⁻³⁴ s), the Higgs-like field is trapped is a state of false vacuum.

The Universe undergoes a super-cooling transition: the temperature decreases below the phase transition point but the Higgs field stays in the false vacuum state.



While the energy density of the Higgs field is positive, the Universe expands at accelerated rate (inflation) and the energy stored in the Higgs field increases.

Inflation stops when the Higgs field decays to the real vacuum.

The energy released by the Higgs field is converted into matter particles.



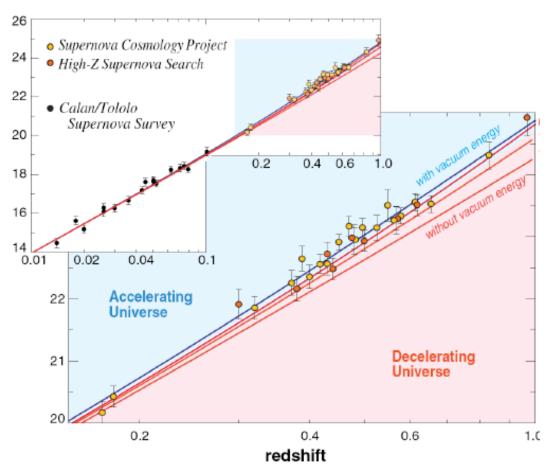
The Universe expansion is accelerating

In 1998, two groups used distant Supernovae to measure the expansion rate of the universe: Perlmutter et al. (Supernova Cosmology Project), and Schmidt et al. (High-z Supernova Team)

They got the same result:

The Universe expansion is accelerating

Some form of energy (dark energy) fills space





Vacuum energy density

Dark energy responsible for acceleration of expansion is very small

From particle physics we know that Vacuum has energy:

- potential energy of scalar fields
- energy of quantum fluctuations as predicted by quantum mechanics

This vacuum energy is 100 orders of magnitude larger than dark energy!

This huge discrepancy is known as the vacuum catastrophe.



The Terascale

The Standard Model would fail at high energy without the Higgs particle or other 'new physics'

Based on the available data and on quite general theoretical insights it was expected that the 'new physics' would manifest at an energy around

1 Tera-electronVolt = 10¹² electronVolt

accessible at the LHC for the first time



End of Lecture 1